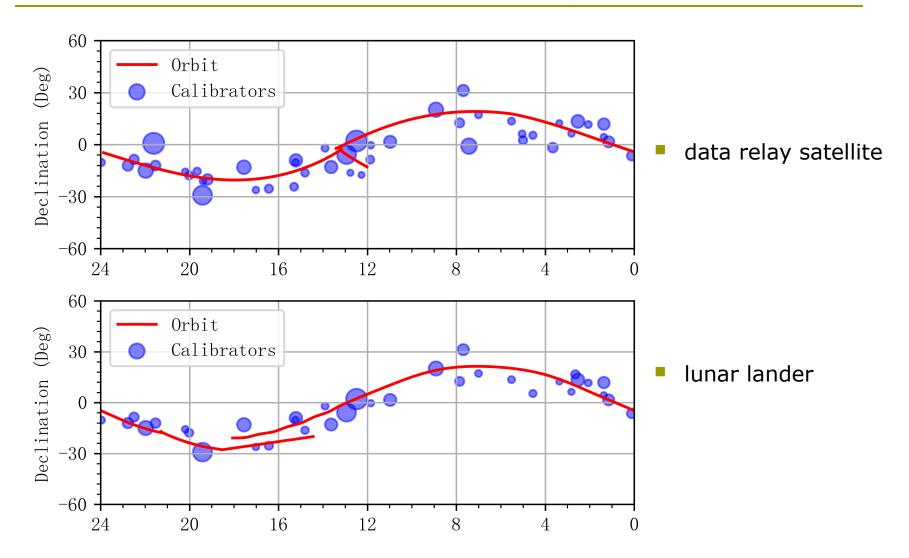
# Contribution of AOV observation ns to the ICRF3

Fengchun Shu

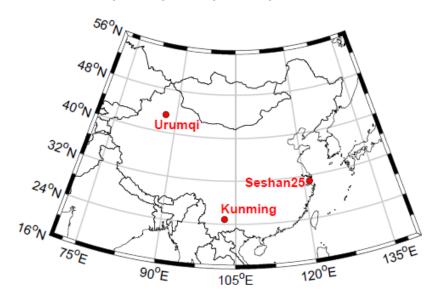
Shanghai Astronomical Observatory

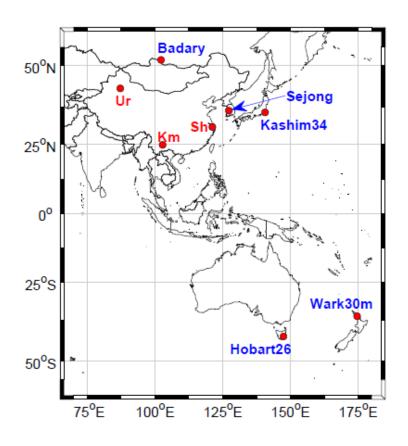
# Calibrator sources for Chang'e-4



# Search for new ecliptic plane sources

- Core stations
  - Sh, Km, Ur
- Cooperation stations
  - Kv, Kb, Ho, Wa, Bd





# Observing sessions in search mode

Date	Codo	Chatiana	#
Date	Code	Stations	#source
yyyy-mm-dd			
2015-02-13	VEPS01	ShKmUr	293
2015-02-14	VEPS02	ShKmUr	338
2015-04-23	VEPS03	ShKmUrKv	300
2015-04-24	VEPS04	ShKmUrKv	400
2015-08-10	VEPS05	ShKmKvHo	252
2015-08-19	VEPS06	ShKmKvHo	277
2016-03-02	VEPS07	ShKmUrKb	333
2016-03-11	VEPS08	ShKmUrKb	477
2016-05-13	VEPS09	ShUrHo	291
2016-05-14	VEPS10	ShUrKv	322
2016-07-06	VEPS11	ShUrKb	307
2016-09-02	VEPS12	ShKmUr	424
2016-09-03	VEPS13	ShKmUr	344
2017-04-15	VEPS14	ShKmUr	408
2017-09-22	VEPS15	ShKmWa	231
2017-09-23	VEPS16	ShKmWa	213
2017-12-14	VEPS17	KmUrBd	291

- Observed sources > 4000
- 2 scans every source
- Total observing time > 400h
- Data volumn > 900TB

#### Note:

Sh-Seshan25,

Km-Kunming,

Ur-Urumqi,

Kv-Sejong,

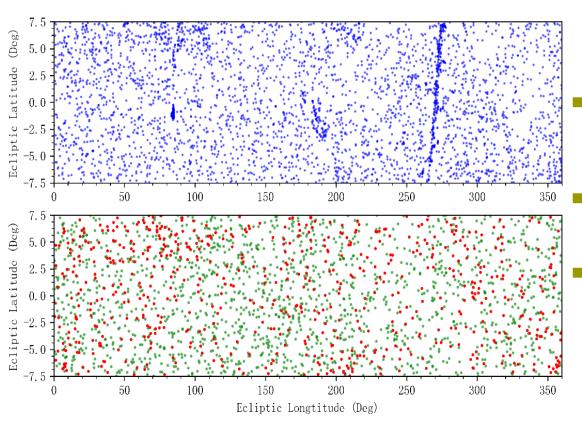
Ho-Hobart26,

Kb-Kashim34,

Wa-Wark30m,

**Bd-Badary** 

# Distribution of ecliptic plane calibrator so urces



- Blue: 3557 sources observed but not detected
  - Red: 662 sources newly detected
- Green: 1554 sources detected in other program

# Absolute astrometry of weak sources

- VLBA
  - BS250: 4\*8h, 110 target sources
  - BS264: 6\*8h, 124 target sources
- EVN
  - ES087(T6+Ef): 24h, 115 sources
- AOV
  - AOV
  - APSG

# AOV sessions scheduled by SHAO

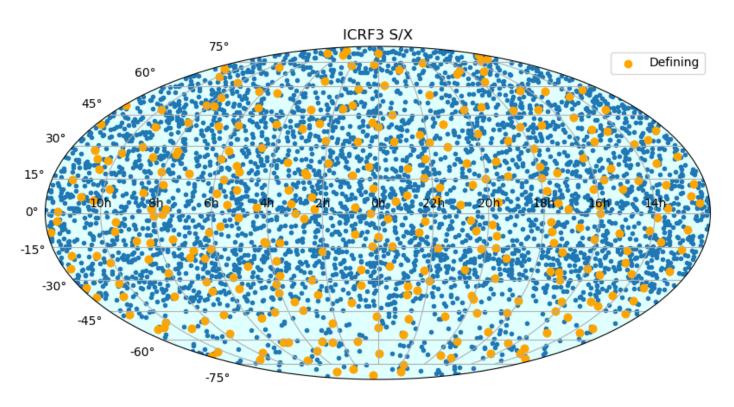
Observing code and number of target sources are shown below

	AOV004(10)						
2016	AOV010(48)	AUA012(38 )	AOV012(27 )	APSG39(20 )			
2017	AOV015(16)	AOV016(41 )	AOV017(70 )	APSG40(16 )	APSG41(16)		
2018	AOV019(78)	AOV021(30 )	AOV025(65 )	AOV027(30 )	AOV029(?)	APSG42(14 )	APSG43(11 )

AOV with Tianma
AOV without Tianma
APSG

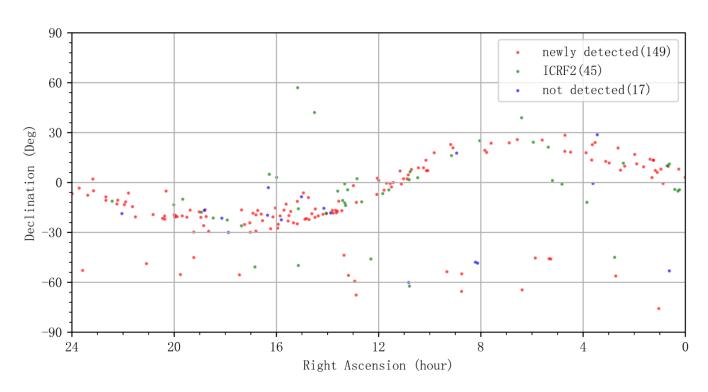
- APSG39-42 were observed at 512Mbps
- All other sessions were observed at 1Gbps

# 4536 S/X sources in the ICRF3 released in 2 018



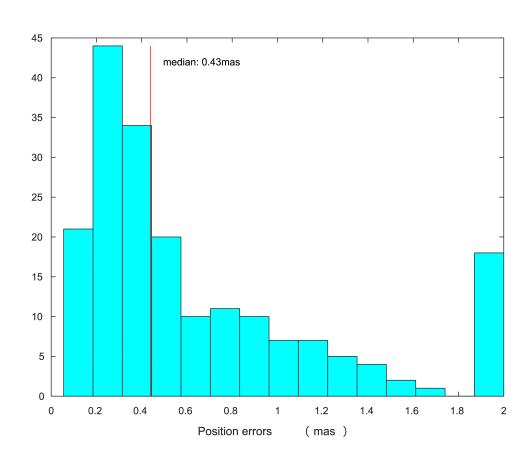
available at http://hpiers.obspm.fr/icrspc/newwww/icrf/index.php

#### Contribution to the ICRF3



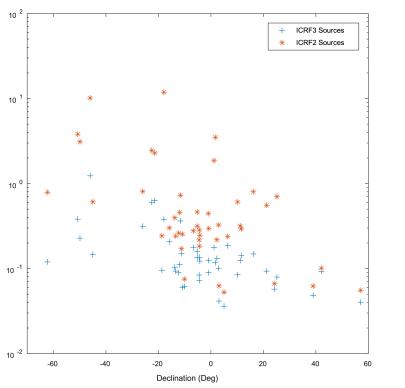
Compared with ICRF2, the ICRF3 has 1122 new sources. Among them, 149 sources (13.3%) have been observed by AOV sessions, and 132 sources were firstly observed in the AOV.

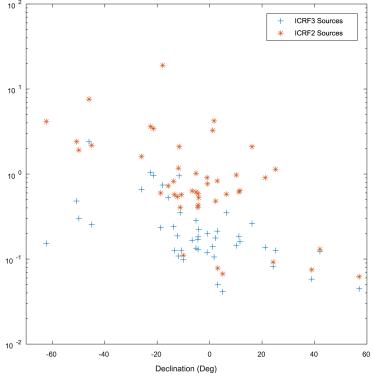
# Statistics on position errors of 194 sources



### Formal errors in RA & Dec

#### 45 AOV sources common to ICRF2 and ICRF3

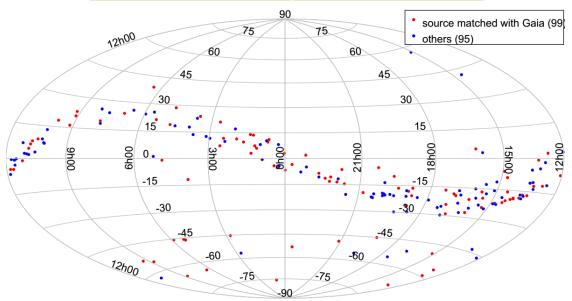




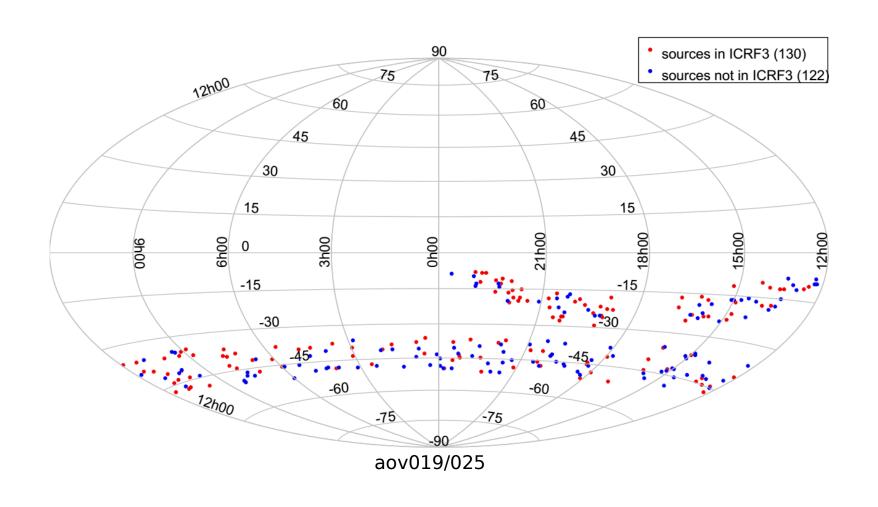
### Cross match with Gaia DR2

#### VLBI/Gaia association for 195 sources

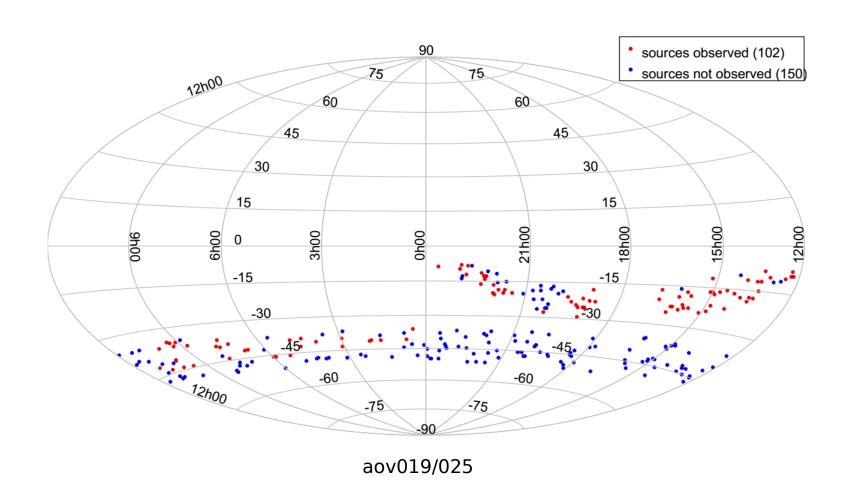
search radius	Matched sources
100 mas	135
50 mas	126
20 mas	99



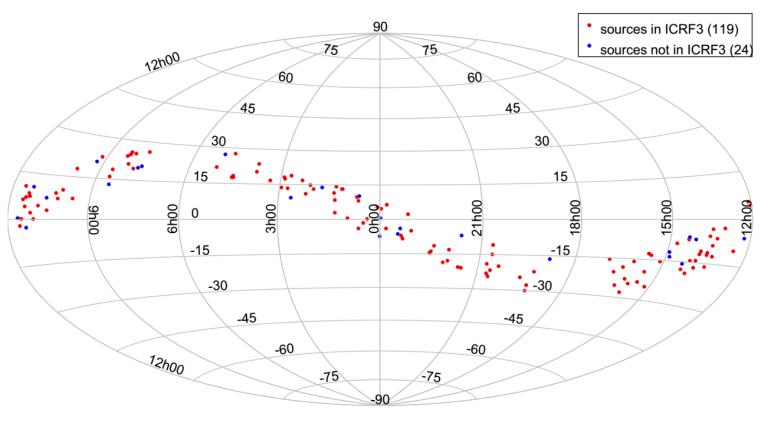
### 2018 AOV sessions with T6



## 2018 AOV sessions with T6

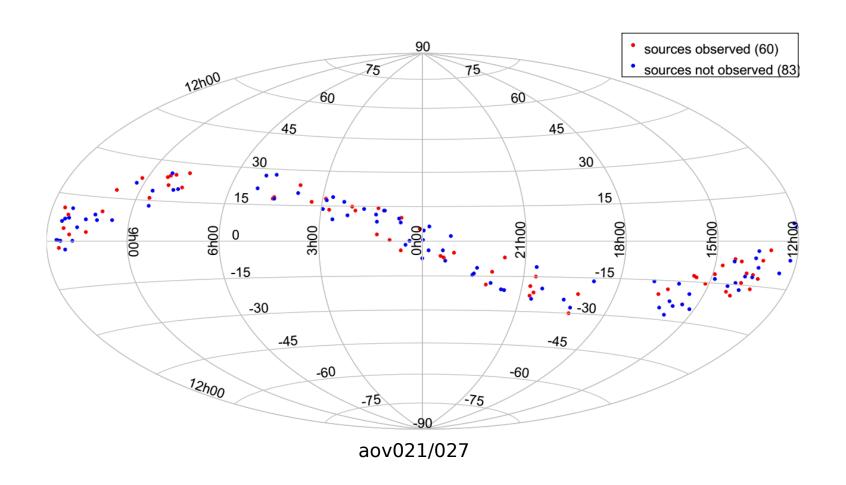


## 2018 AOV sessions without T6



aov021/027

### 2018 AOV sessions without T6



Thanks for your attention.